

# TeV-Scale Sterile Neutrinos

Manfred Lindner



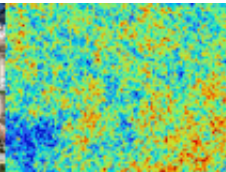
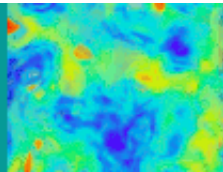
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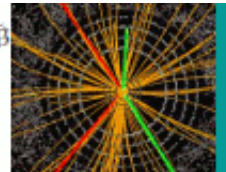
**SHEP**

Southampton  
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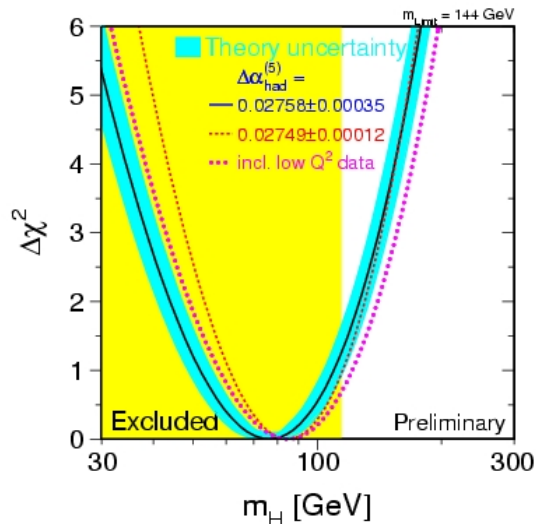
Feb. 08, 2013



$$\begin{aligned} & \Delta_{\text{S}}^{\text{G}} \\ & t - 2m_{\nu}^2 \bar{B}(m_{\nu}^2, t) + \frac{1}{4} \bar{B} \\ & + \Delta_{\text{S}}^{\text{P}} \\ & \sim \bar{B}(m_{\nu}^2, t) \end{aligned}$$

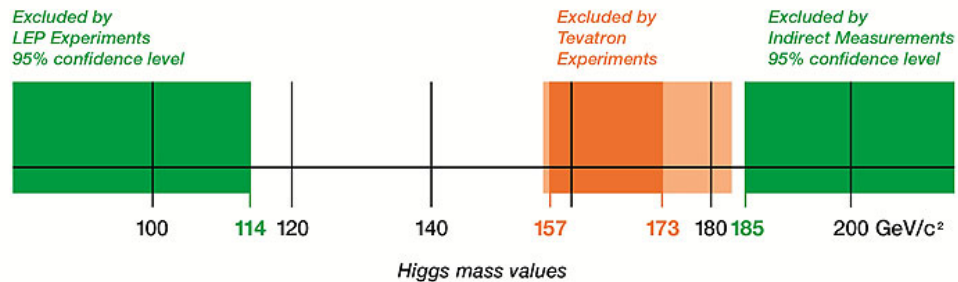


# SM works perfectly & Higgs seems to be there



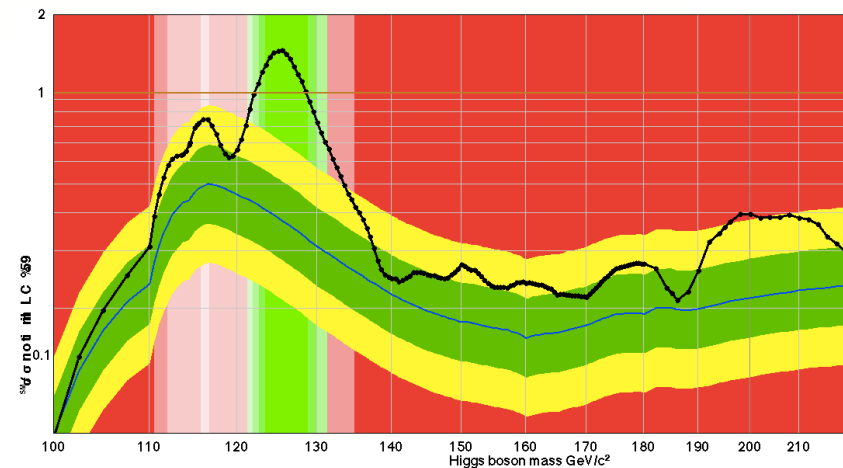
## Search for the Higgs Particle

Status as of March 2011

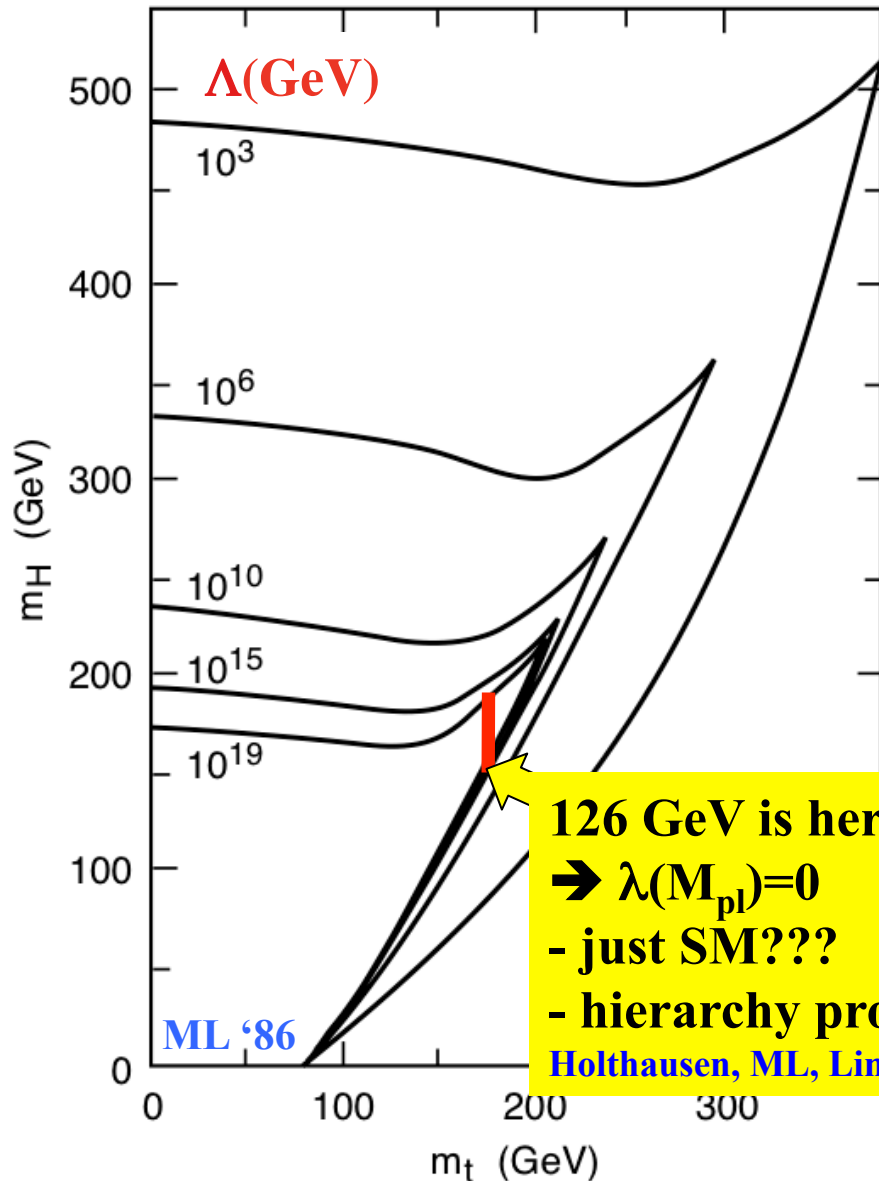


- mass range was shrinking...
- ... and is now rather precisely known
- no signs for anything else
- ➔ just the SM?

- It's a well defined QFT
- just like QED, QCD, ...
  - ➔ 18+1 parameters ➔ experimental input
  - ➔ tests by over-constraining (S,T,U, running  $\alpha$ 's, ...)
  - ➔ per se no hierarchy problem!  $\leftrightarrow$  embedding QFT self-consistency ➔

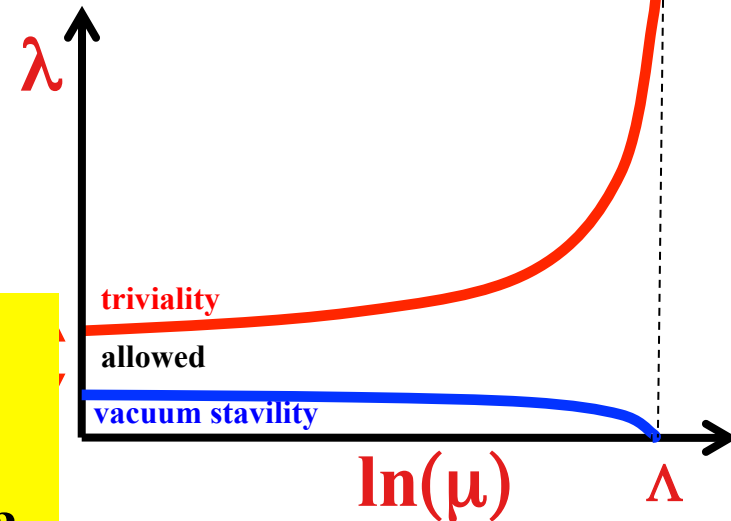


# Triviality and Vacuum Stability



$$126 \text{ GeV} < m_H < 174 \text{ GeV}$$

SM does not exist w/o embedding  
- U(1) coupling, Higgs self-coupling



126 GeV is here!  
→  $\lambda(M_{pl})=0$   
- just SM???  
- hierarchy problem?  
Holthausen, ML, Lim

→ RGE arguments seem to work  
→ we need some embedding

# The SM works perfectly but it must be extended....

- **Many theoretical reasons for BSM physics:**
  - **Flavour problem ...**  $\leftrightarrow$  maybe insights via neutrinos
  - **Hierarchy problem**
    - only a problem once the SM is embedded...  $\rightarrow$  GUTs, ...
    - **separation of two scalar scales:** SM Higgs and some other scalar at  $\Lambda$
    - Planck scale physics  $\rightarrow$  new concepts, spin 2, no new scalars?
      - $\rightarrow$  no 2<sup>nd</sup> scalar for direct embeddings at the Planck scale?
      - $\rightarrow$  SM boundary conditions that point to  $\Lambda_{\text{Planck}}$   $\rightarrow$  vacuum stability line!
- **Experimental facts:**
  - **SM cannot explain Baryon Asymmetry of the Universe**
    - BUT: neutrino masses require SM extension  $\rightarrow$  **SM+** $\rightarrow$
    - leptogenesis = one of the best BAU explanations
  - **Dark Matter**
    - some extra particle is required which is DM
    - particles connected to the hierarchy problem, strong CP, ...  $\rightarrow$  neutrinos
  - **Neutrino Masses and maybe evidences for sterile neutrinos**

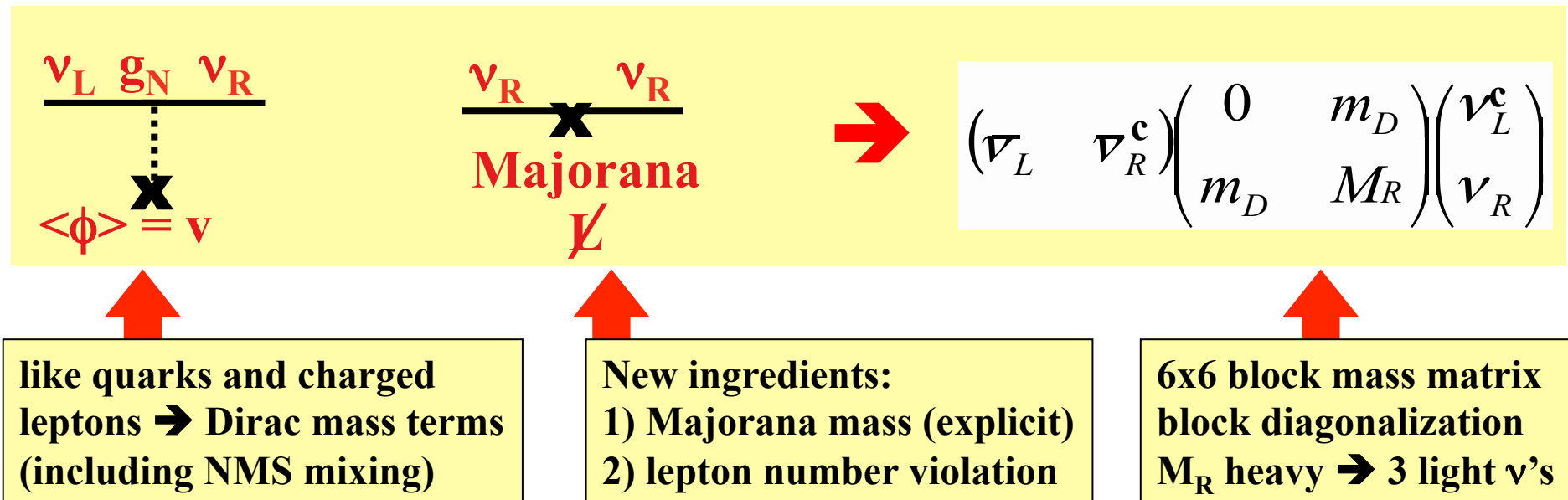
# New Physics: Neutrino Mass Terms

SM  $\sim m\phi\bar{L}R = (2,1)$

→ new fields

→ Simplest possibility:  
add 3 right handed neutrino fields

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$
$L_Q = \begin{pmatrix} l_u \\ l_d \end{pmatrix}$	3	2	1/3
$r_u$	3	1	4/3
$r_d$	3	1	-2/3
$L_L = \begin{pmatrix} l_\nu \\ l_e \end{pmatrix}$	1	2	-1
$r_\nu ???$	1	1	0
$r_e$	1	1	-2



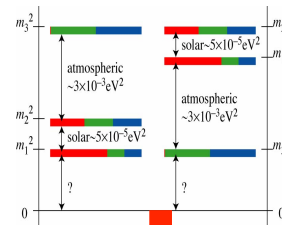
**NEW ingredients, 9 parameters → SM+ and sea-saw**

# Suggestive Seesaw Features

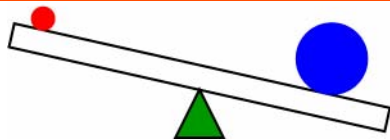
QFT: natural value of mass operators  $\leftrightarrow$  scale of symmetry

$m_D \sim$  electro-weak scale

$M_R \sim$  L violation scale  $\leftarrow ? \rightarrow$  embedding (GUTs, ...)



See-saw (type I)



$$m_\nu = m_D M_R^{-1} m_D^T$$

$$m_h = M_R$$

?  $\rightarrow$  EW scale

Numerical hints:

For  $m_3 \sim (\Delta m_{\text{atm}}^2)^{1/2}$ ,  $m_D \sim$  leptons  $\rightarrow M_R \sim 10^{11} - 10^{16} \text{ GeV}$   
 $\rightarrow$   $\nu$ 's are **Majorana particles**,  $m_\nu$  probes  $\sim$  GUT scale physics!  
 $\rightarrow$  smallness of  $m_\nu \leftrightarrow$  high scale of  $L$ , symmetries of  $m_D, M_R$

# The Neutrino Spectrum

## The standard picture:

3 heavy sterile neutrinos typ.  $\geq 10^{13}$  GeV

→ leptogenesis, role in GUTs, ...

Some mechanism which makes  
1, 2, ... heavy states light?

→ light sterile neutrino(s)

→ tiny heavy-light mixing expected  
 $\theta^2 < O(m_\nu/m_s)$

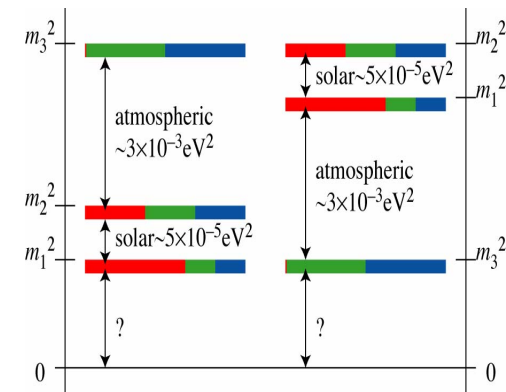
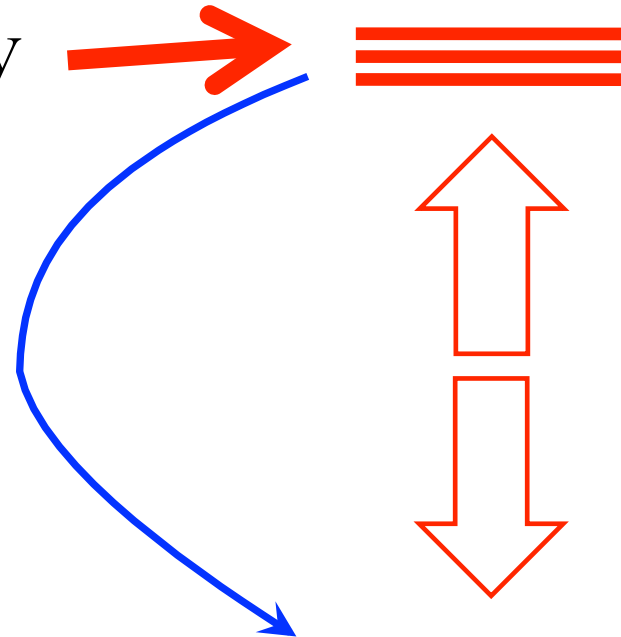
3 light active neutrinos

→ this could easily be wrong

- more than 3  $N_R$  states, ...

-  $M_R$  may have special eigenvalues, ...

→ light sterile neutrinos ?!



# Evidences for Light Sterile Neutrinos

## Particle Physics:

Reactor anomaly, LSND, MiniBooNE, MINOS, Gallex...

- New and better data / experiments are needed to clarify the situation
- maybe something exciting around the corner?
- would hint at eV scale and sizable percentage type mixings

CMB: Extra eV-ish neutrinos possible J. Hamann et al. , ...

BBN: Extra  $\nu$ 's possible:  $N_\nu \simeq 3.7 \pm 1$

E. Aver, K. Olive, E. Skillman (2010), Y. Izotov, T. Thuan(2010)

Astrophysics: keV-ish sterile neutrinos could explain pulsar kicks

Kusenko, Segre, Mocioiu, Pascoli, Fuller et al., Biermann & Kusenko, Stasielak et al., Loewenstein et al., Dodelson, Widrow, Dolgov, ...

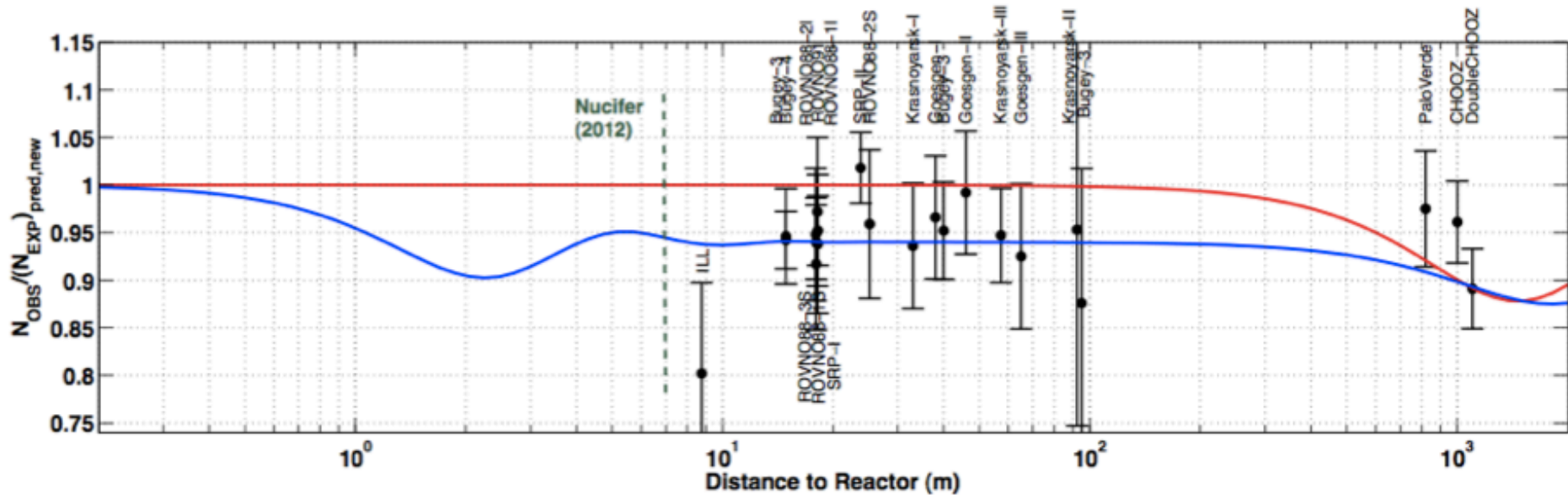
My believe: Most likely not all of them are correct

→ but any one has far reaching consequences!



# The Reactor Anomaly

New reactor fluxes and global reactor data: Mention et al.



$\theta_{13}$  can reduce flux at  $L \geq 1$  km, but not at shorter baselines  
Sterile neutrino with  $\Delta m^2 \sim \text{eV}$  can nicely account for reduction

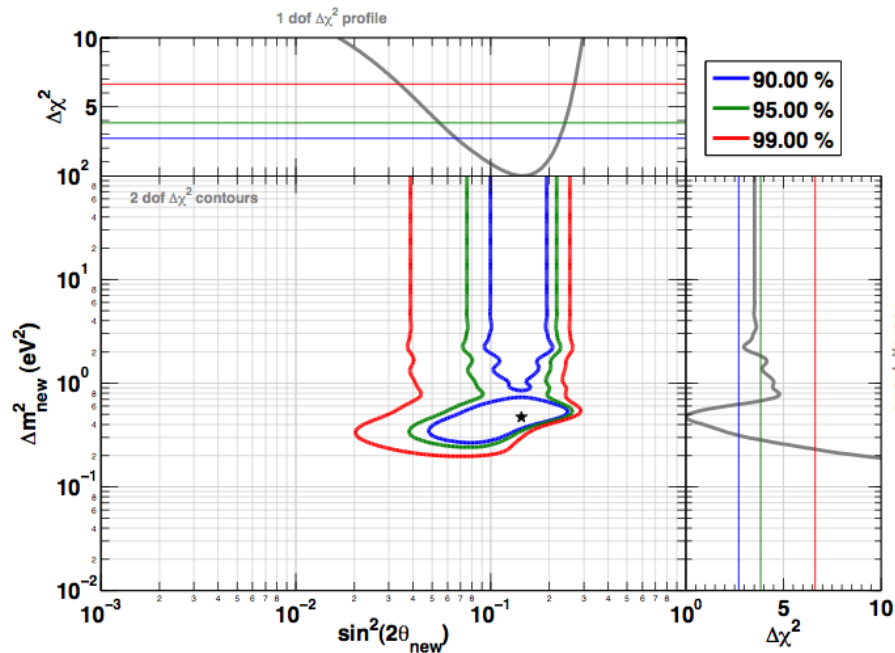
→ 3+1 fits ; all evidences for eV scale → 3+2, ...

See e.g. T. Schwetz at NEUTRINO 2012

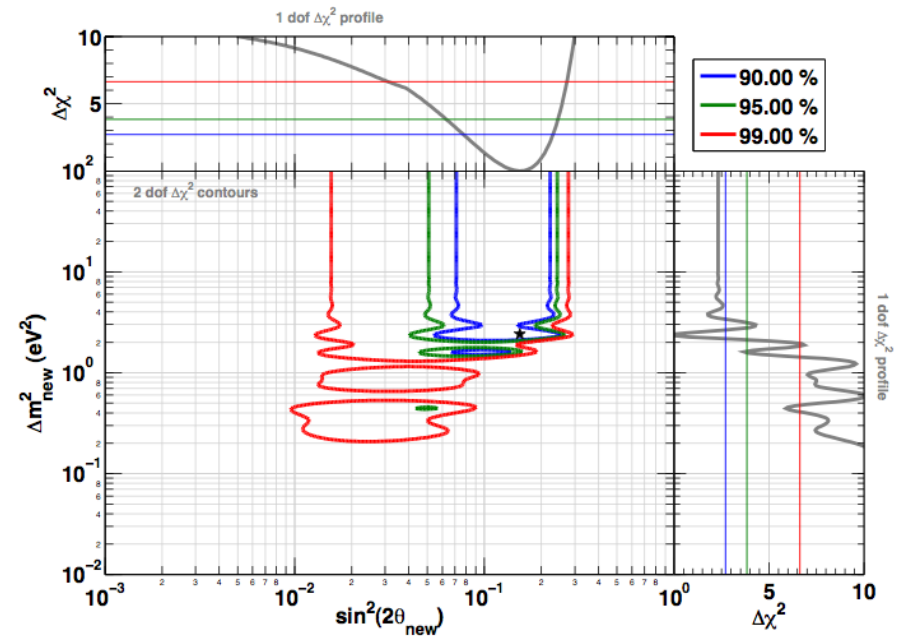
→ will be tested by new experiments (e.g. NUCIFER @ few meters)

# Reactor Anomaly

Rate only analysis



Rate + Shape (Bugey 3) analysis



- Best fit:  $\Delta m^2 = 2.4 \text{ eV}^2$  ;  $\sin^2(2\theta_{\text{new}}) = 0.14$
- 2.9 sigma significance

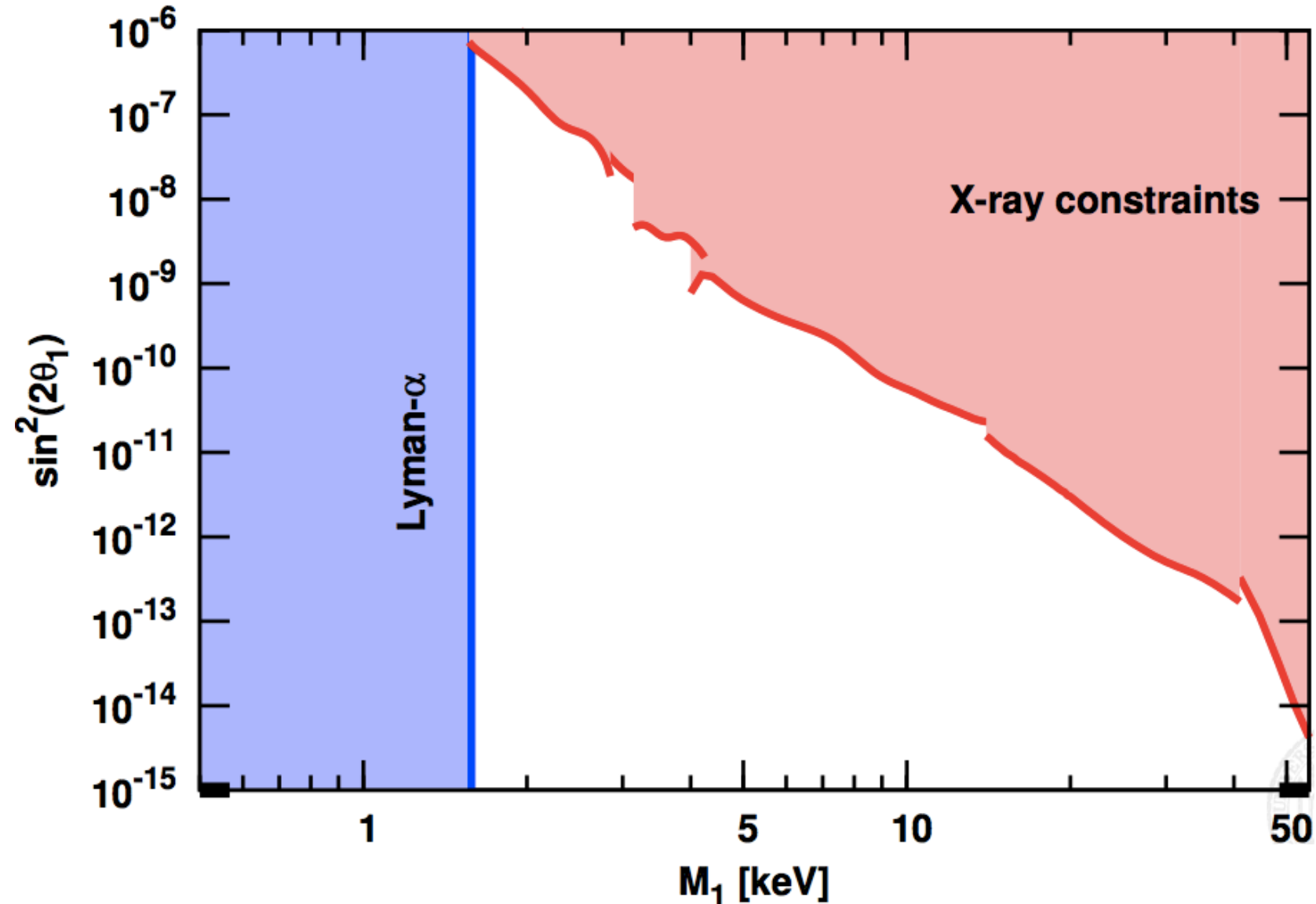
# Could Neutrinos be Dark Matter?

- Active neutrinos would be perfect Hot Dark Matter → ruled out:
  - destroys small scale structures in cosmological evolution
  - measured neutrino masses too small → maybe HDM component
- keV sterile neutrinos: Warm Dark Matter → workes very well:
  - relativistic at decoupling
  - non-relativistic at radiation to matter dominance transition
  - OK for  $M_X \simeq \text{few keV}$  with very tiny mixing
  - reduced small scale structure → smoother profile, less dwarf satellites
  - scenario where one sterile neutrino is keV-ish, the others heavy
  - tiny active – sterile mixings  $O(m_\nu/M_R)$

**Note: Right-handed neutrinos exist probably anyway – just make one light!**

Asaka, Blanchet, Shaposhnikov, Asaka, Shaposhnikov; Kusenko, Segre, Mocioiu, Pascoli, Fuller et al., Biermann & Kusenko, Stasielak et al., Loewenstein et al., Dodelson, Widrow, Dolgov, ... Bezrukov, Hettmannsperger, ML

# Allowed Range for keV Sterile $\nu$ Dark Matter



# Could TeV-Scale Sterile Neutrinos Exist?

**E. Akhmedov, A. Kartavtsev, ML, L. Michaels and J. Smirnov**  
**arXiv 1302.1872**

Let's assume that TeV-scale sterile neutrinos exist.

- what are the limits?
- do they improve things or make them worse?

**→ improvements of over-all fits!**

# Simplest Extension of the SM

- Add 3 right-handed neutrinos (Majorana fields)  $N_i = N_i^c$

$$\mathcal{L} = \mathcal{L}_{SM} + \frac{1}{2}\bar{N}_i(i\cancel{\partial} - M_i)N_i - h_{\alpha i}\bar{\ell}_\alpha\tilde{\phi}N_i - h_{i\alpha}^\dagger\bar{N}_i\tilde{\phi}^\dagger\ell_\alpha$$

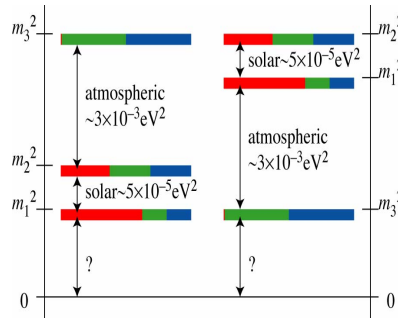
- This can solve experimentally required facts :
  - small neutrino masses via type I see-saw
  - baryon asymmetry of the Universe via leptogenesis
  - eV-ish sterile neutrinos (reactor anomaly)
  - keV sterile neutrinos (dark matter)
  - ...
- Explaining the sterile spectrum ...???
- Study effects of sterile neutrinos with any mass  $\rightarrow$  TeV

# N=3 sterile Neutrinos

sterile neutrinos  $\nu_{4,5,6}$   
spectrum ???



active neutrinos  $\nu_{1,2,3}$



6x6 mixing matrix:

$$U = \begin{pmatrix} \mathcal{U} & \mathcal{R} \\ \mathcal{W} & \mathcal{V} \end{pmatrix}$$

(3 × 3) PMNS matrix  $\mathcal{U}$   
is not exactly unitary  
→ deviations

$$\epsilon_\alpha \equiv \sum_{i \geq 4} |U_{\alpha i}|^2$$

$$\epsilon_e - \epsilon_\mu = 0.0022 \pm 0.0025$$

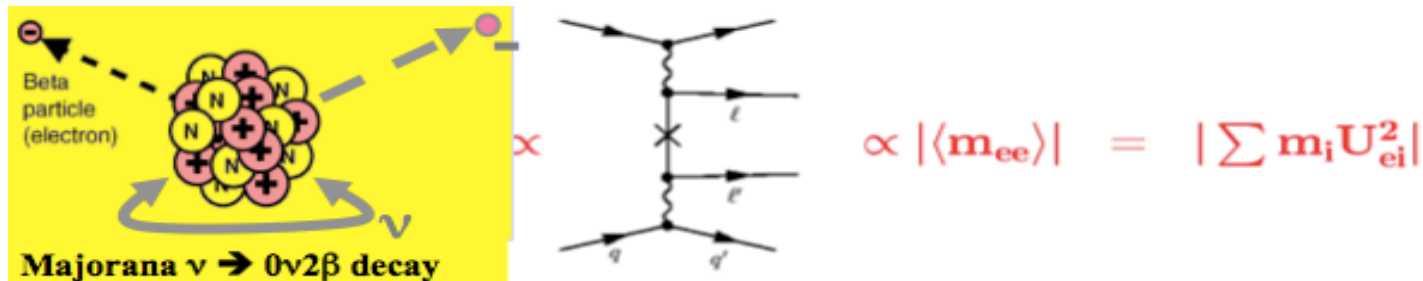
$$\epsilon_\mu - \epsilon_\tau = 0.0017 \pm 0.0038$$

$$\epsilon_e - \epsilon_\tau = 0.0039 \pm 0.0040$$

Global fits of active neutrino data  
usually assume 3x3 unitarity!  
→ unconstrained fits  $\simeq$  up to 4%  
changes Antusch et al.

# Consequences of sterile Neutrinos

- Shifts in active neutrino parameters in global analyses
- L-violating admixture in light  $\nu$ 's  $\rightarrow 0\nu\beta\beta$  Beta Decay



- **Effective mass including all states:**

$$|\langle m_{ee} \rangle| \approx \left| \sum_{i=1}^3 U_{ei}^2 m_i - \sum_{i=4}^{3+n} F(A, M_i) U_{ei}^2 m_i \right|$$

$$F(A, m_i) \approx (m_a/m_i)^2 f(A)$$

$f(A)$  depends on the decaying isotope

today  $|\langle m_{ee} \rangle| < 0.4 \text{ eV} \rightarrow$  will improve in next years!



- **Additional interactions of W and Z:**

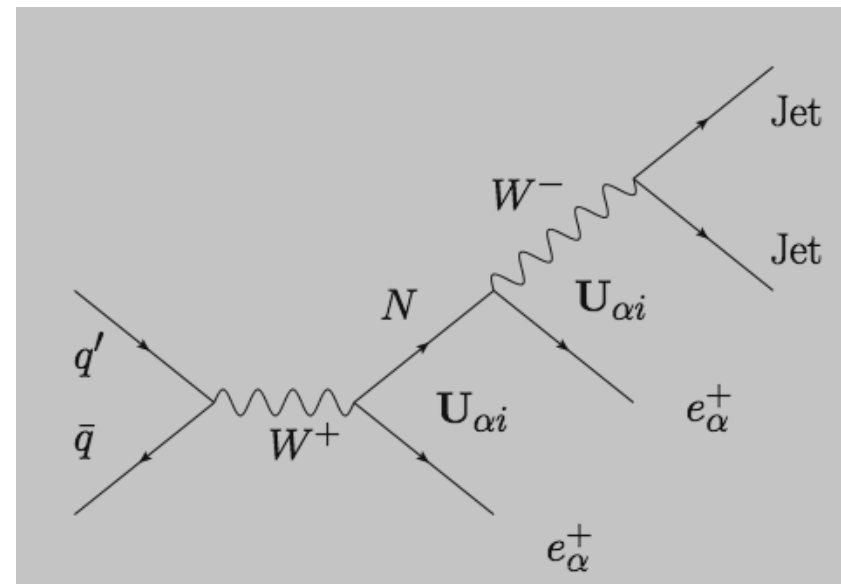
$$\mathcal{L}_{\text{int}} = -\frac{e}{2c_w s_w} Z_\mu \sum_{i,j=1}^{3+n} \sum_{\alpha=e,\mu,\tau} \bar{\nu}_i \mathbf{U}_{i\alpha}^\dagger \gamma^\mu P_L \mathbf{U}_{\alpha j} \nu_j$$

$$- \frac{e}{\sqrt{2}s_w} W_\mu \sum_{i=1}^{3+n} \sum_{\alpha=e,\mu,\tau} \bar{\nu}_i \mathbf{U}_{i\alpha}^\dagger \gamma^\mu P_L e_\alpha + \text{h.c.}$$

→ **L-violating processes**

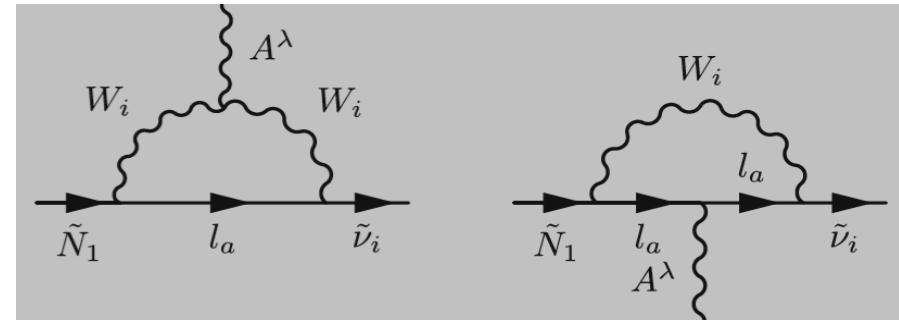
e.g. **production of heavy neutrino at LHC as intermediate state**

→ **after upgrade to 14TeV with  $100 \text{ fb}^{-1}$  expected limit**



$|\sum_i \mathbf{U}_{\alpha i}^2 m_i^{-1}| \simeq 6.5 \cdot 10^{-3} \text{ TeV}^{-1}$  **sensitivity to 800 GeV sterile neutrinos!**

- **L-violating decays;**  
e.g. :  $\mu \rightarrow e\gamma$



$$\text{BR}(\mu \rightarrow e\gamma) = \frac{\Gamma(\mu \rightarrow e\gamma)}{\Gamma(\mu \rightarrow e\nu\bar{\nu})} = \frac{3\alpha}{32\pi} |\delta_\nu|^2$$

$$\delta_\nu = 2 \sum_i \mathbf{U}_{ei}^* \mathbf{U}_{\mu i} g(m_i^2/M_W^2)$$

$$g(x) = \int_0^1 \frac{(1-\alpha)d\alpha}{(1-\alpha) + \alpha x} [2(1-\alpha)(2-\alpha) + \alpha(1+\alpha)x].$$

$$\rightarrow \delta_\nu = 2 \sum_{i=4}^{3+n} \mathbf{U}_{ei}^* \mathbf{U}_{\mu i} [g(m_i^2/M_W^2) - 5/3]$$

**MEG bound:**  $\text{BR}(\mu^+ \rightarrow e^+\gamma) \leq 2.4 \cdot 10^{-12}$

- **Non-unitarity in neutrino oscillations**  
→ zero distance effect in addition to oscillations

$$P_{\alpha\beta}(L = 0) = \frac{\delta_{\alpha\beta} (1 - 2\epsilon_{\alpha}) + \epsilon_{\alpha}\epsilon_{\beta}}{(1 - \epsilon_{\alpha})(1 - \epsilon_{\beta})}$$

- not easy to detect  $\leftrightarrow$  normalization
- note that this cannot explain the reactor anomaly!

- **Electro-weak precision observables**
  - tree level mixing and loop effects (S,T,U)
  - may explain some 'deviations'
  - fix SM definition → predict other observables

EWPO	Theory (Standard Model)	Experiment
$\Gamma_{\text{lept}}$ (MeV)	$84.005 \pm 0.015$	$83.984 \pm 0.086$
$\Gamma_{\text{inv}}/\Gamma_{\text{lept}}$	$5.9721 \pm 0.0002$	$5.942 \pm 0.016$
$\sin^2 \theta_W$	$0.23150 \pm 0.0001$	$0.2324 \pm 0.0012$
$g_L^2$	$0.3040 \pm 0.0002$	$0.3026 \pm 0.0012$
$g_R^2$	$0.0300 \pm 0.0002$	$0.0303 \pm 0.0010$
$M_W$ (GeV)	$80.359 \pm 0.011$	$80.385 \pm 0.015$

- **Z-width very precisely known  $\simeq 2\sigma$  too low**
- **NuTeV anomaly**

- **Sterile neutrinos lead to modifications**  
     **→ may explain some 'deviations'**

$$\Gamma_{\text{inv}} / [\Gamma_{\text{inv}}]_{\text{SM}} = \frac{1}{3} \sum_{\alpha} (1 - \epsilon_{\alpha})^2$$

**CC and NC scattering  $\leftrightarrow$  mixing**

$$\sigma_{\alpha}^{\text{CC}} = \sigma_{\alpha, \text{SM}}^{\text{CC}} (1 - \epsilon_{\alpha}),$$

$$\sigma_{\alpha}^{\text{NC}} = \sigma_{\alpha, \text{SM}}^{\text{NC}} (1 - \epsilon_{\alpha})^2$$

**definition of  $G_{\mu}$ :**

$$G_{\mu}^2 = G_F^2 (1 - \epsilon_{\mu})(1 - \epsilon_e)$$

## Corrections due to heavy sterile $\nu$ 's:

- Both tree level ( $\epsilon$ 's) and loop (S,T,U) effects
- U is tiny
- identical or almost identical parameter combinations
- cancellations
- some observables don't change
- some do  $\leftrightarrow$  data

$$\frac{\Gamma_{\text{lept}}}{[\Gamma_{\text{lept}}]_{\text{SM}}} = 1 + 0.6 (\epsilon_e + \epsilon_\mu + 0.0145 T) - 0.0021 S,$$

$$\frac{\Gamma_{\text{inv}}/\Gamma_{\text{lept}}}{[\Gamma_{\text{inv}}/\Gamma_{\text{lept}}]_{\text{SM}}} = 1 - 0.67 (\epsilon_e + \epsilon_\mu + \epsilon_\tau) + 0.0021 S - 0.0015 T,$$

$$\frac{\sin^2 \theta_w^{\text{lept}}}{[\sin^2 \theta_w^{\text{lept}}]_{\text{SM}}} = 1 - 0.72 (\epsilon_e + \epsilon_\mu + 0.0145 T) + 0.0016 S,$$

$$\frac{g_L^2}{[g_L^2]_{\text{SM}}} = 1 + 0.41 \epsilon_e - 0.59 \epsilon_\mu - 0.0090 S + 0.0022 T,$$

$$\frac{g_R^2}{[g_R^2]_{\text{SM}}} = 1 - 1.4 \epsilon_e - 2.4 \epsilon_\mu + 0.031 S - 0.0067 T,$$

$$\frac{M_W}{[M_W]_{\text{SM}}} = 1 + 0.11 \epsilon_e + 0.11 \epsilon_\mu - 0.0036 S + 0.0056 T + 0.0042 U$$

- **Details...**

$$\begin{aligned}
S_{\text{tot}} = S_N + S_{\text{SM}} = & -\frac{1}{2\pi M_Z^2} \\
& \times \left[ \sum_{i,j=1}^{3+n} \sum_{\alpha\beta} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha j} \mathbf{U}_{j\beta}^\dagger \mathbf{U}_{\beta i} \Delta Q(M_Z^2, m_i^2, m_j^2) \right. \\
& + \sum_{i,j=1}^{3+n} \sum_{\alpha\beta} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha j} \mathbf{U}_{i\beta}^\dagger \mathbf{U}_{\beta j} m_i m_j \Delta B_0(M_Z^2, m_i^2, m_j^2) \\
& + \sum_{\alpha} m_{\alpha}^2 B_0(0, m_{\alpha}^2, m_{\alpha}^2) + \sum_{\alpha} Q(M_Z^2, m_{\alpha}^2, m_{\alpha}^2) \\
& \left. - 2 \sum_{\alpha} m_{\alpha}^2 B_0(M_Z^2, m_{\alpha}^2, m_{\alpha}^2) \right], \quad (20)
\end{aligned}$$

$$\begin{aligned}
Q(q^2, m_1^2, m_2^2) \equiv & (D-2)B_{22}(q^2, m_1^2, m_2^2) \\
& + q^2 [B_1(q^2, m_1^2, m_2^2) + B_{21}(q^2, m_1^2, m_2^2)]
\end{aligned}$$

$B_0, B_1, B_{21}$  and  $B_{22}$

**are the usual loop functions**

$$\begin{aligned}
T_{\text{tot}} = T_N + T_{\text{SM}} = & -\frac{1}{8\pi s_w^2 M_W^2} \\
& \times \left[ \sum_{i,j=1}^{3+n} \sum_{\alpha\beta} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha j} \mathbf{U}_{j\beta}^\dagger \mathbf{U}_{\beta i} Q(0, m_i^2, m_j^2) \right. \\
& + \sum_{i,j=1}^{3+n} \sum_{\alpha\beta} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha j} \mathbf{U}_{i\beta}^\dagger \mathbf{U}_{\beta j} m_i m_j B_0(0, m_i^2, m_j^2) \\
& - 2 \sum_{i=1}^{3+n} \sum_{\alpha} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha i} Q(0, m_i^2, m_{\alpha}^2) \\
& \left. + \sum_{\alpha} m_{\alpha}^2 B_0(0, m_{\alpha}^2, m_{\alpha}^2) \right], \quad (1)
\end{aligned}$$

$$\begin{aligned}
U_{\text{tot}} = U_N + U_{\text{SM}} = & \frac{1}{2\pi M_Z^2} \\
& \times \left[ \sum_{i,j=1}^{3+n} \sum_{\alpha\beta} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha j} \mathbf{U}_{j\beta}^\dagger \mathbf{U}_{\beta i} \Delta Q(M_Z^2, m_i^2, m_j^2) \right. \\
& + \sum_{i,j=1}^{3+n} \sum_{\alpha\beta} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha j} \mathbf{U}_{i\beta}^\dagger \mathbf{U}_{\beta j} m_i m_j \Delta B_0(M_Z^2, m_i^2, m_j^2) \\
& + \sum_{\alpha} m_{\alpha}^2 B_0(0, m_{\alpha}^2, m_{\alpha}^2) - \sum_{\alpha} Q(M_Z^2, m_{\alpha}^2, m_{\alpha}^2) \\
& - 2 \sum_{\alpha} m_{\alpha}^2 B_0(M_Z^2, m_{\alpha}^2, m_{\alpha}^2) \\
& \left. - 2(M_Z/M_W)^2 \sum_{\alpha} \mathbf{U}_{i\alpha}^\dagger \mathbf{U}_{\alpha i} \Delta Q(M_W^2, m_i^2, m_{\alpha}^2) \right]. \quad (21)
\end{aligned}$$

# Results of global Fits to all Data

- Assume Lagrangian including see-saw type I  
 $\leftrightarrow$  parameter relations
- All data: LFV, LHC, EWPO and active neutrinos  
 and consider 3 typical mass spectra

$$\sin^2 \theta_{12} = 0.30 \pm 0.013,$$

$$\sin^2 \theta_{23} = 0.41_{-0.025}^{+0.037},$$

$$\sin^2 \theta_{13} = 0.023 \pm 0.0023,$$

$$\delta_{CP} = 300_{-138}^{+66},$$

	NH	IH	QD
$m_1$ (eV)	$\sim 0$	$4.85 \cdot 10^{-2}$	$\sim 0.1$
$m_2$ (eV)	$8.660 \cdot 10^{-3}$	$4.93 \cdot 10^{-2}$	$\sim 0.1$
$m_3$ (eV)	$4.97 \cdot 10^{-2}$	$\sim 0$	$\sim 0.1$

- **Quality of fit:**

$$\chi_{\text{EWPO}}^2 = \sum_i \frac{(O_i - O_{i,\text{SM}})^2}{(\delta O_i)^2 + (\delta O_{i,\text{SM}})^2}$$

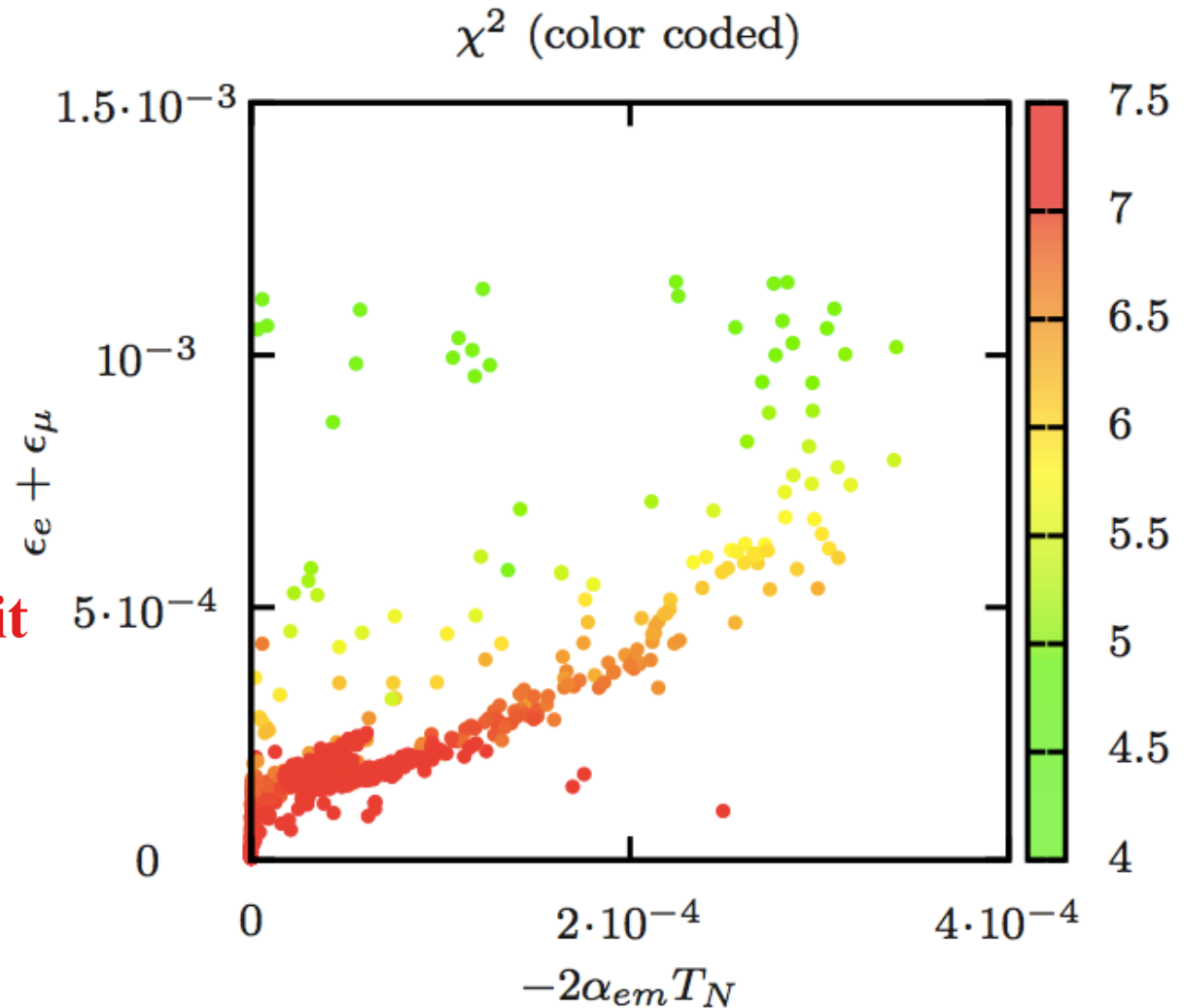


# Some Examples

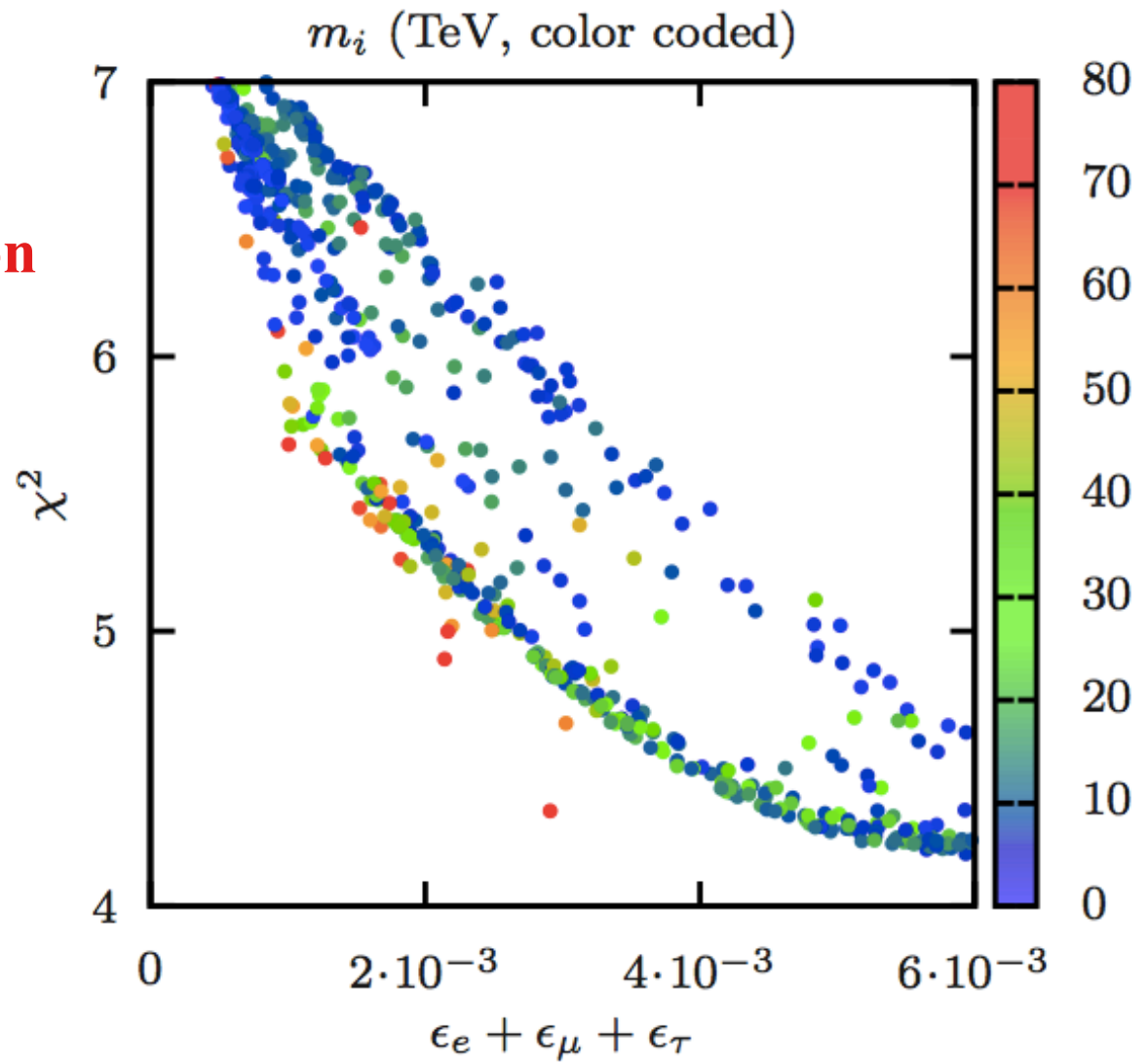
NH,  $\chi^2$  for 4 d.o.f.  
fit to all EWPOs

→ overall  $\chi^2/\text{dof}$  goes  
from  $\simeq 2$  to  $\simeq 1$   
for  $\epsilon_e + \epsilon_\mu \sim 10^{-3}$

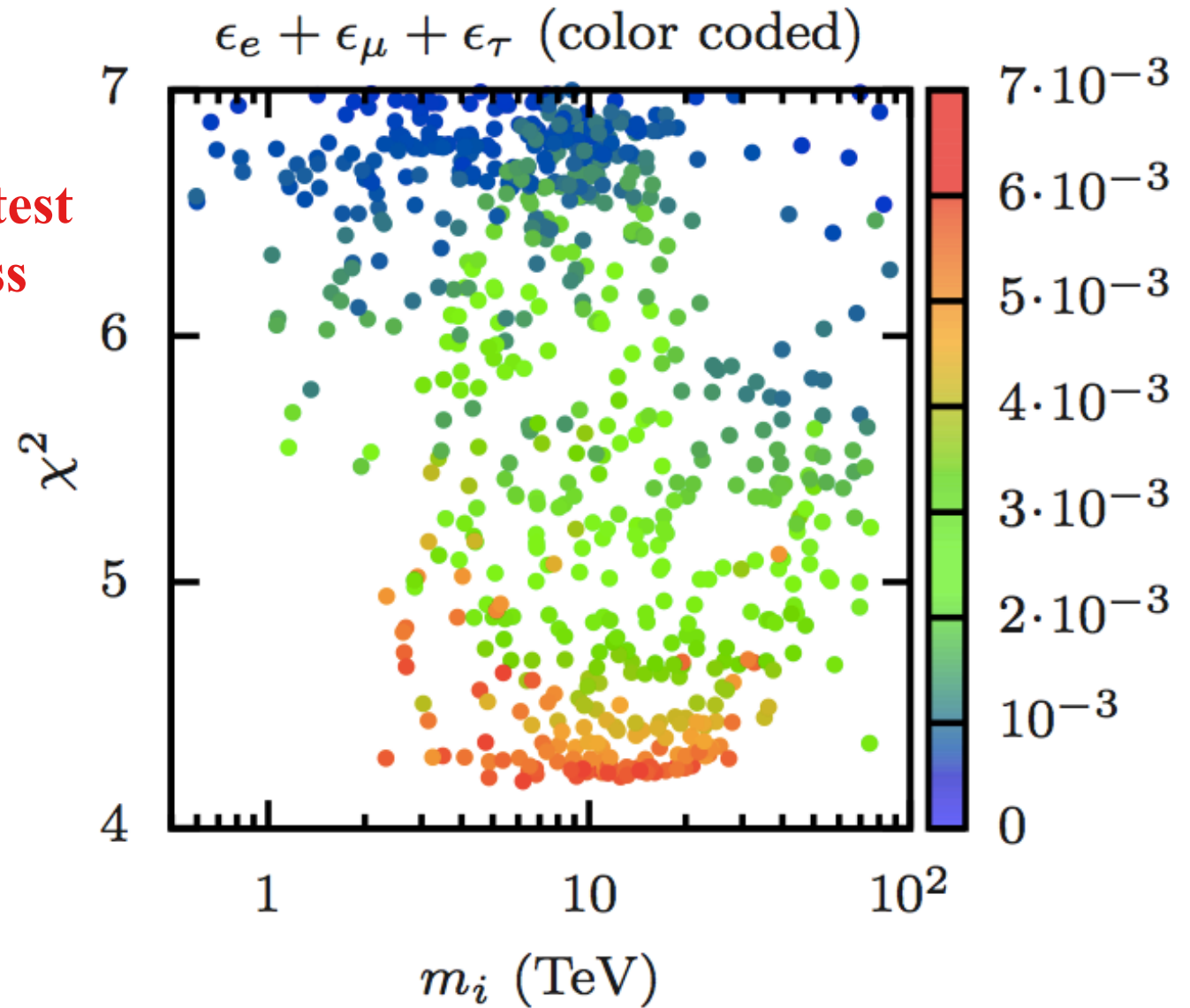
→ sterile  $\nu$ 's improve fit



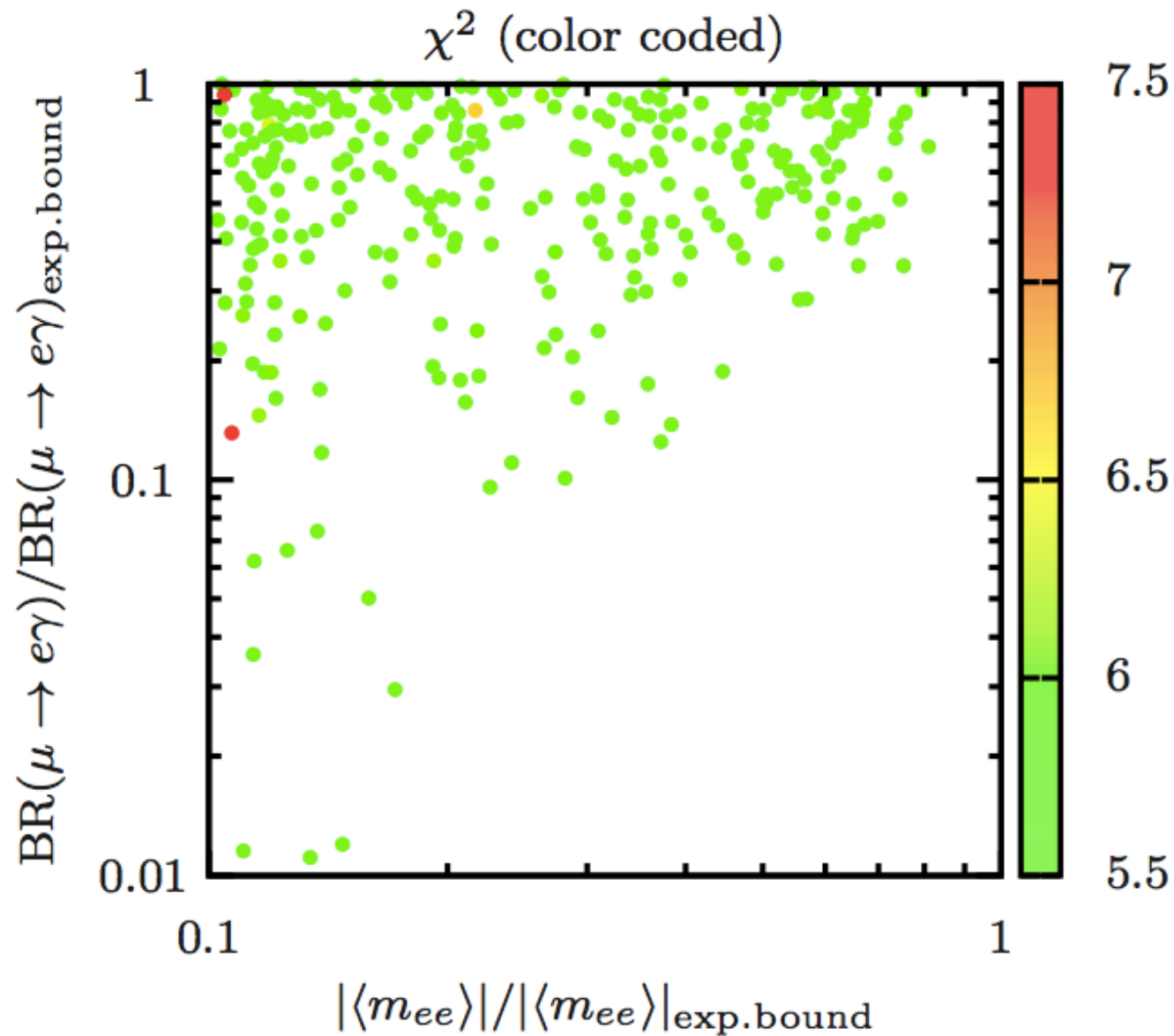
NH, the lightest heavy  
neutrino mass as function  
of  $\chi^2$  and  $\epsilon_e + \epsilon_\mu + \epsilon_\tau$   
for 4dof



**NH,  $\epsilon_e + \epsilon_\mu + \epsilon_\tau$  as a function of the lightest heavy neutrino mass for 4dof**



**IH,  
double beta decay  
and  $\mu \rightarrow e\gamma$  limits  
for 4dof**

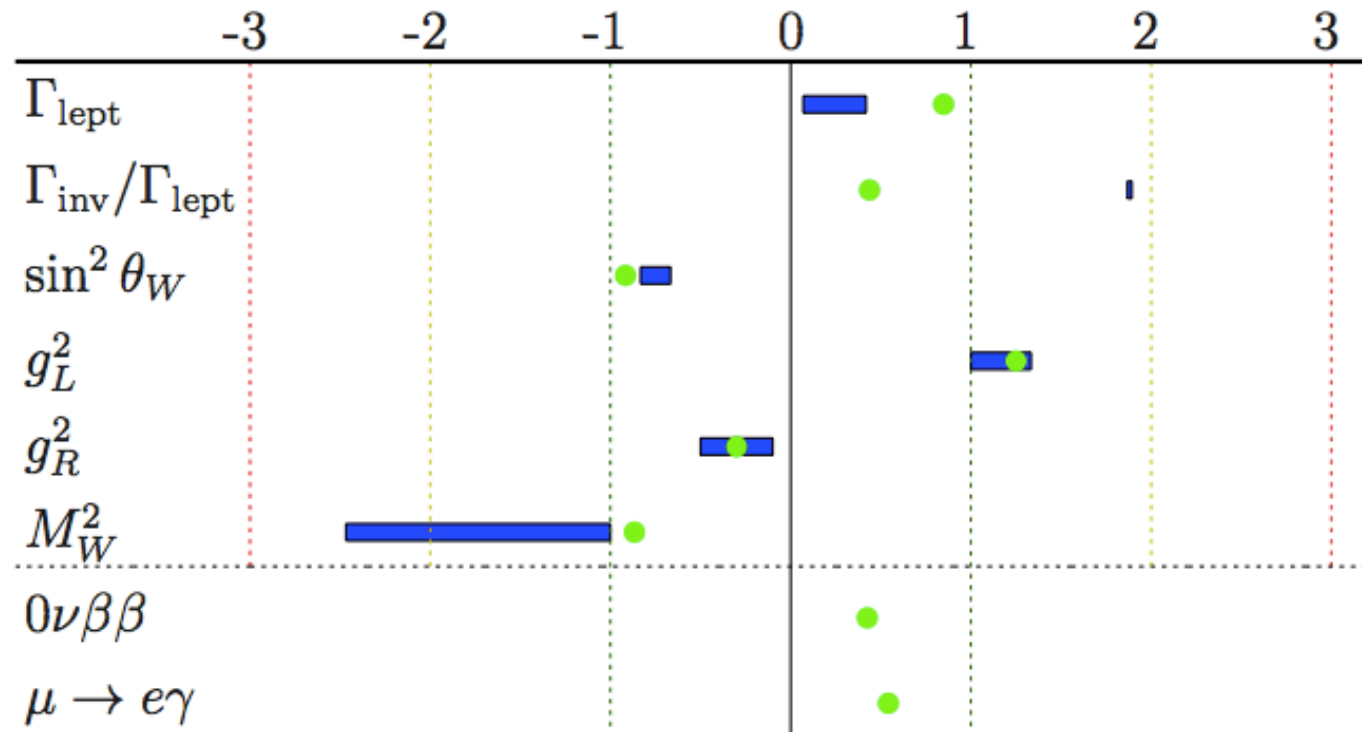


# Which Improvements

Agreement with data = 0 + 1, 2, 3  $\sigma$  ; assumed scenario: NH  
 SM best fit  $\rightarrow$  blue bars

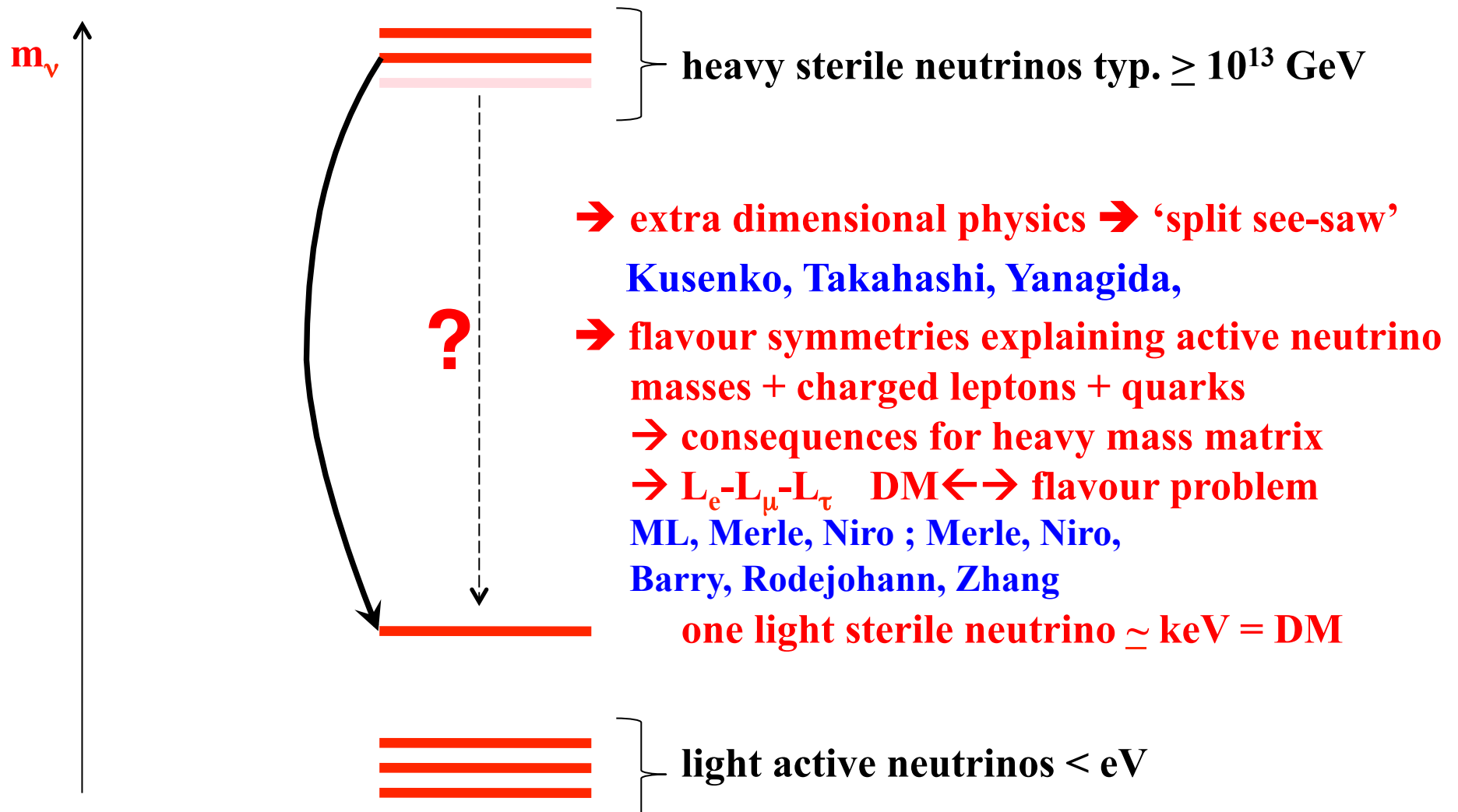
best fit with sterile n's  $\rightarrow$  green points

$M_1 = 20.3$  TeV,  $M_2 = 14.1$  TeV,  $M_3 = 21.0$  TeV,  
 $\varepsilon_e = 2.1 \cdot 10^{-3}$ ,  $\varepsilon_\mu = 3.0 \cdot 10^{-6}$  and  $\varepsilon_\tau = 4.5 \cdot 10^{-3}$



# Explaining keV-ish Sterile Neutrinos

Possible scenario: See-saw + a reason why 1 sterile  $\nu$  is light



# Conclusions

- **Different indications / motivations for sterile neutrinos**  
→ what about masses  $O(\text{TeV})$ ?
- **Interesting cancellations between tree and loop effects of heavy sterile neutrinos**
- **Heavy sterile neutrinos improve global fits to all data and reduces tensions: deviation of Z-width, NuTeV, global**
- **Mass range from a few hundred GeV to 100 TeV**
- **$O(10^{-3})$  mixings**
- **Natural theoretical explanation of the spectrum required**  
→ eV and keV states might also be explained  
→ interesting possibilities / combinations